

Freeform imaging from Hadamard matrices

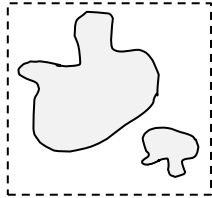
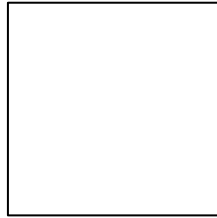
Nicolas Ducros^{1, 2}

¹CREATIS, Univ Lyon, INSA-Lyon, UCB Lyon 1, CNRS, Inserm, CREATIS UMR 5220, U1206, Lyon, France

²IUF, Institut Universitaire de France

Joint work with J Cohen and L Mahieu-Williame

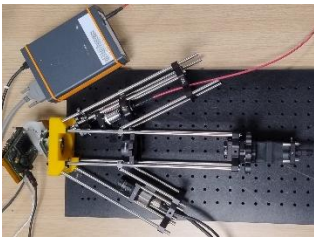
1. Introduction



2. Freeform imaging

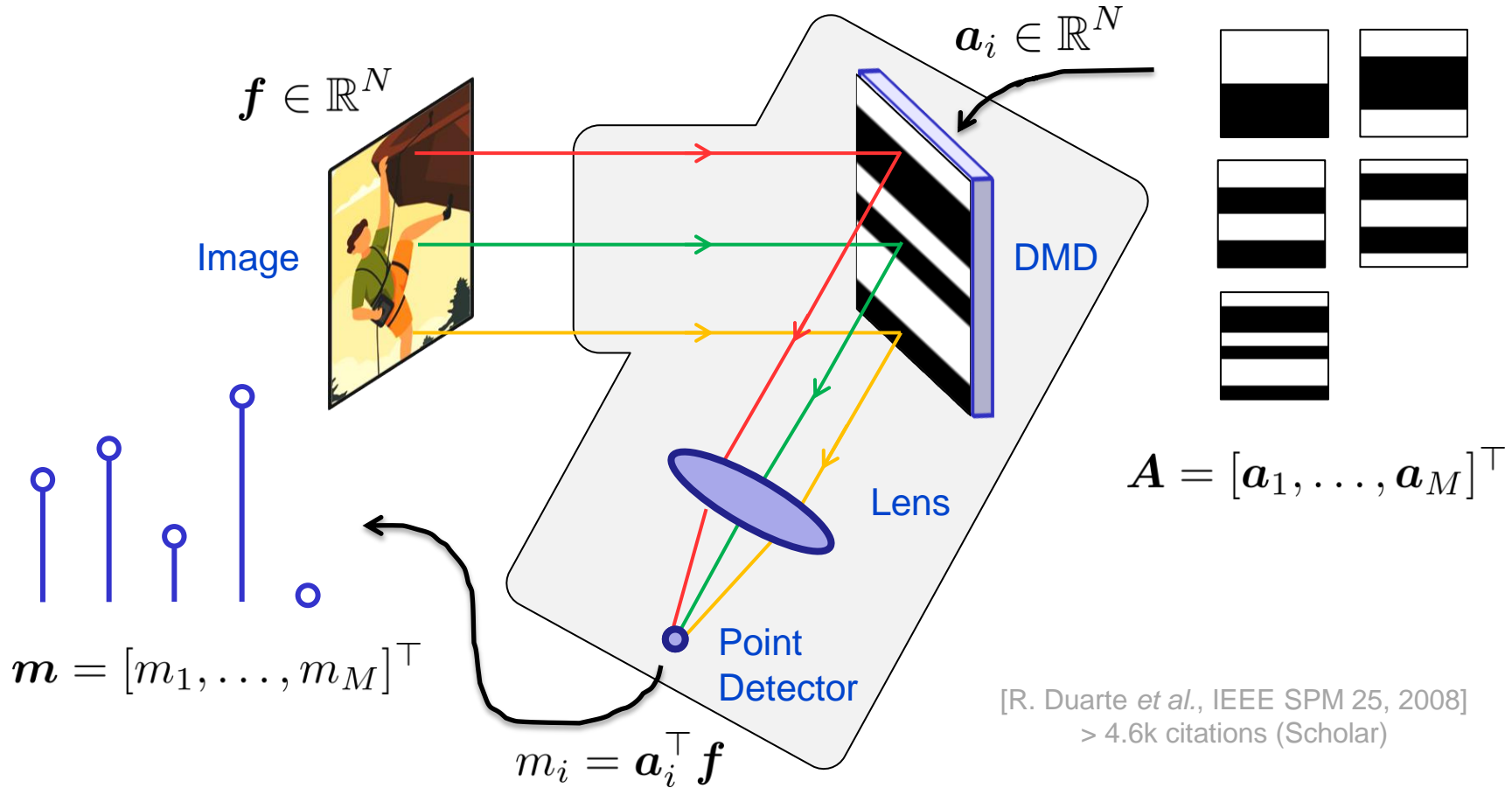
3. Multiplex (or Fellgett's) advantage

$$\text{Mse}(\mathbf{f}^*) \approx \frac{1}{t} [2Q\bar{f} + 4Mf_{\text{ref}}]$$



4. Experimental results

1. Introduction



Acquisition

$$m = Af$$

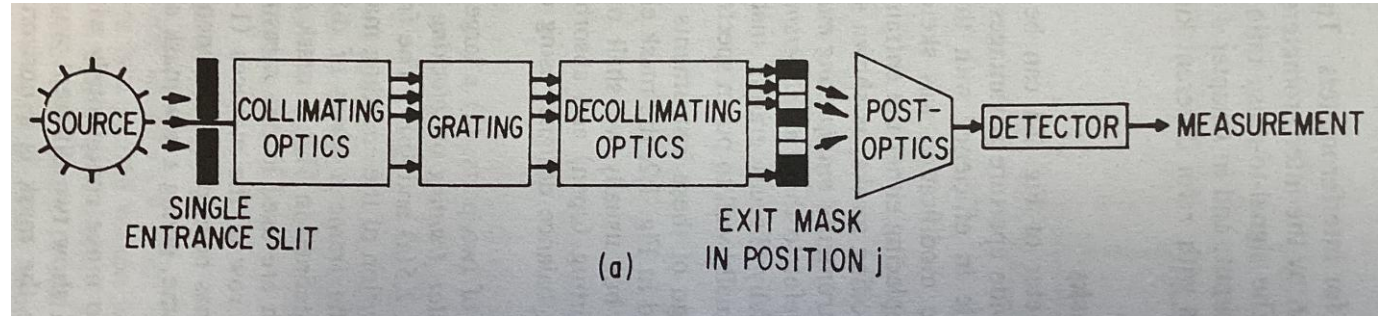
$$A \in \mathbb{R}^{M \times N}$$

Reconstruction ($M \ll N$)

$$f^* \leftarrow \min_f \|m - Af\|_2^2 + \lambda \|\Phi f\|_1$$

HADAMARD TRANSFORM OPTICS

Martin Harwit
Neil J.A. Sloane

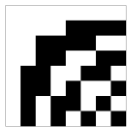


[M. Harwit and N
Sloane, Academic
Press, 1979]

*'(...) conventional spectrometer is modified by using a
mask to encode the light at the output'*

Hadamard

$$\mathbf{A} = \mathbf{H} \in \{-1, 1\}^{N \times N}$$



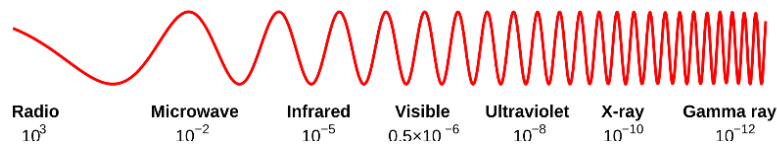
$$\mathbf{H}^T \mathbf{H} = N \mathbf{I}_N$$

L2 reconstruction

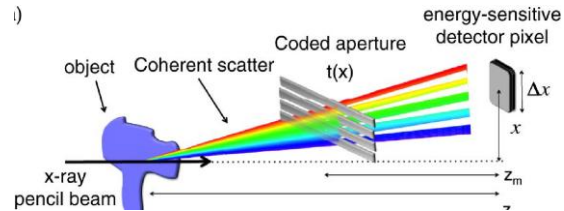
$$\mathbf{f}^* = \frac{1}{N} \mathbf{H}^T \mathbf{m}$$

Fellgett's advantage

$$\text{var}(f_n^*) = \frac{1}{N} \sigma^2 < \sigma^2$$



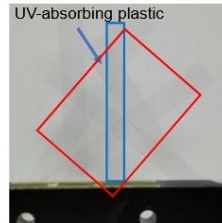
X rays



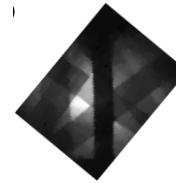
[J. Greenberg *et al.*, Optics letters 39, 2014]

Ultraviolet

visible

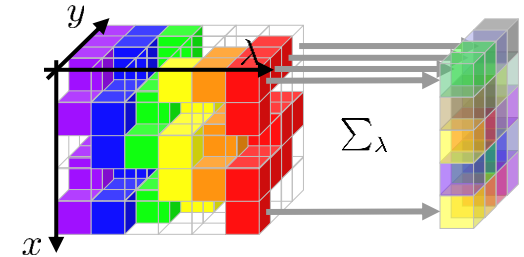


UV



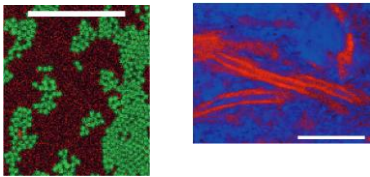
[J. T. Ye *et al.*, Appl. Phys. Lett. 123, 2023]

Hyperspectral



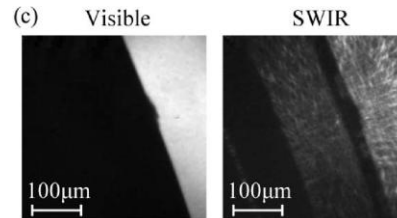
[G Beneti Martins *et al.*, Opt. Express, 2023]
[E Hemsley *et al.*, JOSAA 37 (12), 2020]

Raman imaging



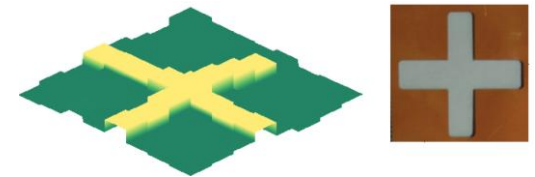
[F Soldevilla *et al.*, Optica 6(3), 2019]
[Scotte *et al.*, Jphys Photonics 5(3), 2023]

Infrared



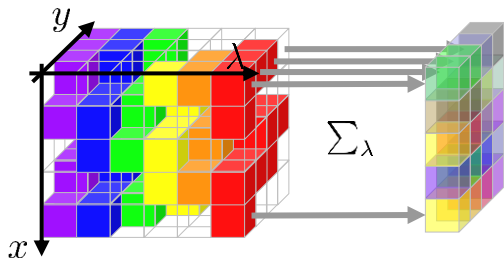
[N. Radwell *et al.*, Optica 1(5), 2014]

Terahertz imaging



[C. Watts *et al.*, Nature Photonics 8(8), 2014]

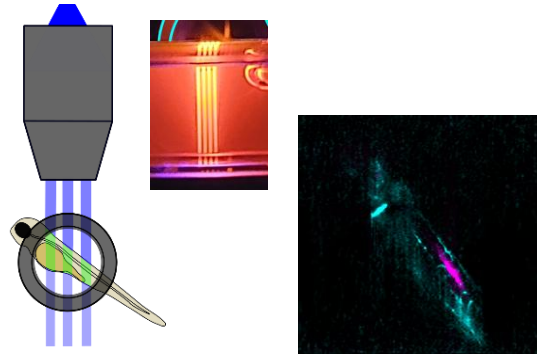
Hyperspectral



200+ hypercubes in open access
<https://pilot-warehouse.creatis.insa-lyon.fr/>

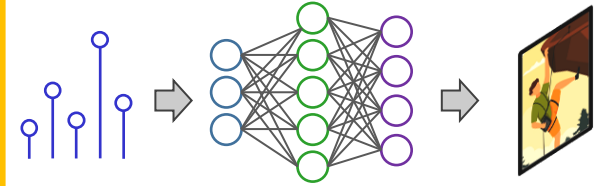
[G Beneti Martins *et al.*, Opt. Express, 2023]

Fluorescence light sheet



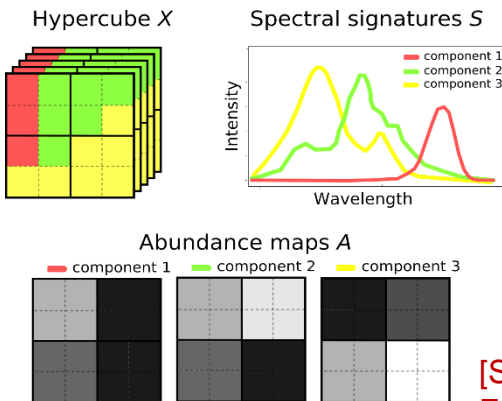
[S Crombez *et al.*, Opt. Express, 2025]

DL-based reconstruction



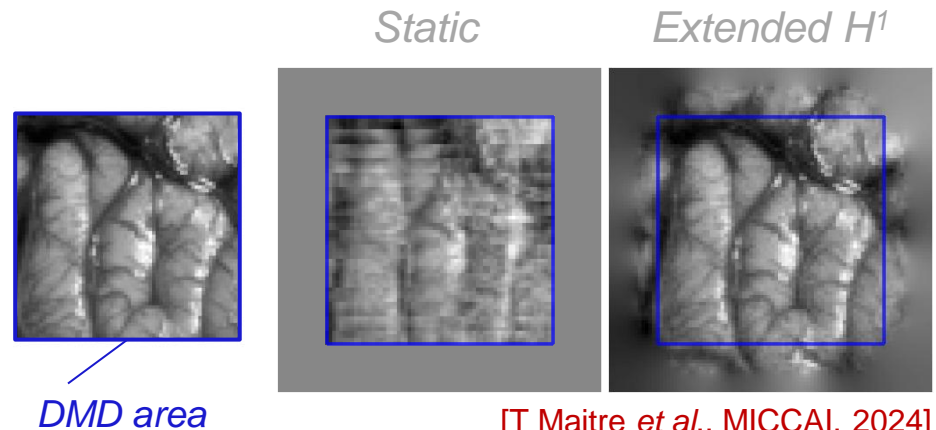
[JFJ. Abascal *et al.*, Opt. Express, 2025]

Spectral unmixing



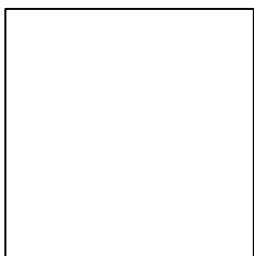
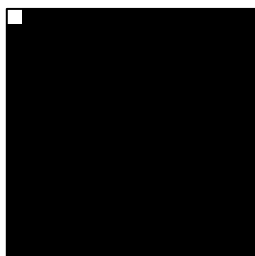
[S Hariga *et al.*, EUSIPCO, 2024]

Motion-compensated imaging

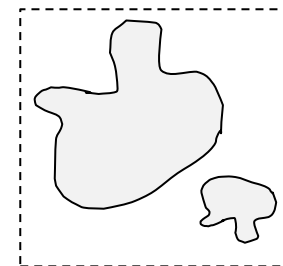
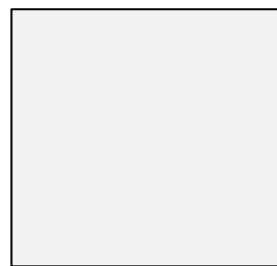


[T Maitre *et al.*, MICCAI, 2024]

Idea#1: Hadamard modulation to reduce noise

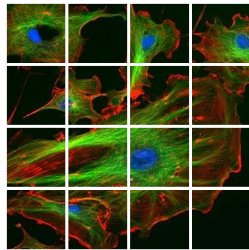


Idea#2: Freeform to capture only relevant information



Easily combined to leverage both advantages

2. Freeform Imaging

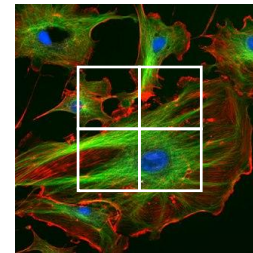


4x4 image

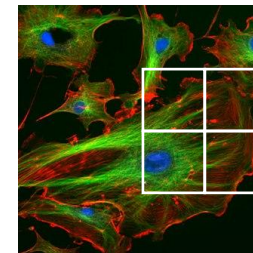
Given a time budget



Fewer pixels =
reduced noise

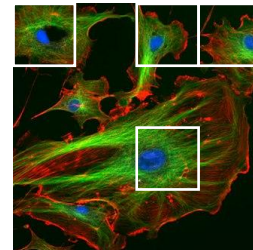


2x2 image



2x2 image

Why not?



4-pixel
image

→ Freeform imaging

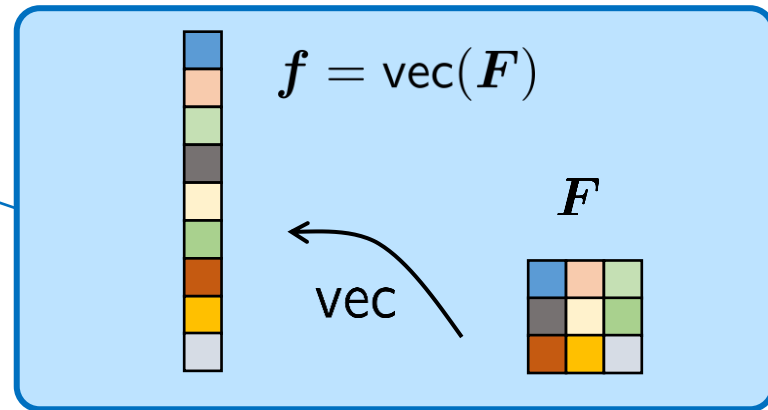
= Capturing an arbitrary
pixel subset within the
FOV.

<https://commons.wikimedia.org/wiki/File:FluorescentCells.jpg>

$$m = Af$$

P : Total number of pixel
within the full FOV

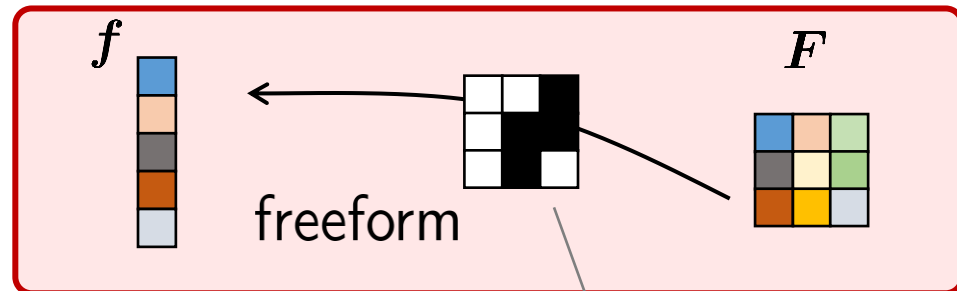
Full (P pixels) Here: $P = 9$ pixels



N : Number of pixels within
the freeform region

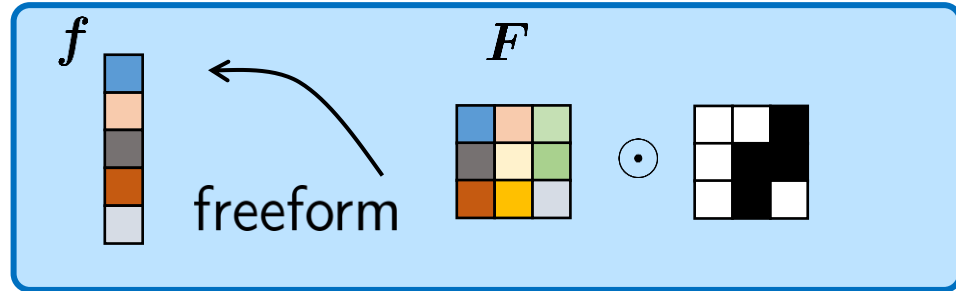
Freeform (N pixels)

Here: $N = 5$ pixels

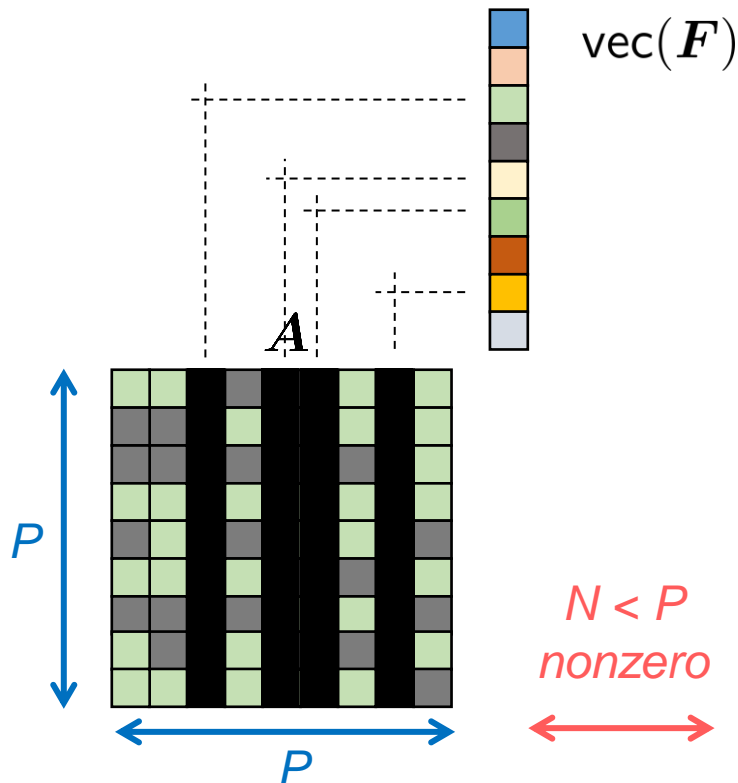


*Mask indicating the pixels
within the freeform region
(white = in; black = out)*

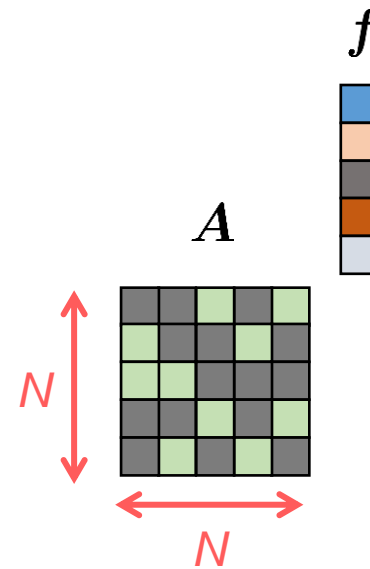
$$m = Af$$



1. Masked matrices



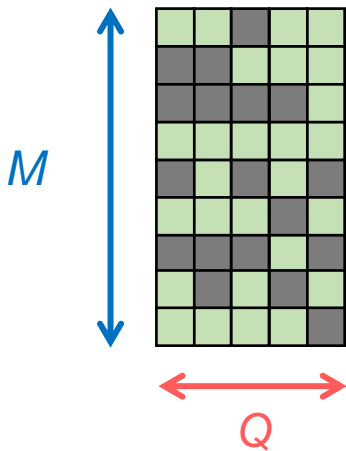
2. Non imaging matrices



3.

Fellgett's advantage in freeform imaging

Acquisition matrix $M \geq Q$



M measurements
for Q pixels

$$\mathbf{A}^T \mathbf{A} = K \mathbf{I}_Q$$

Mean Squared Error* (MSE)

$$\text{Mse}(\mathbf{f}^*) \approx \frac{1}{t} [2Q\bar{f} + 4Mf_{\text{ref}}]$$

$M \geq Q$

Reference flux
(depends only on
experimental parameters)

Number of pixels

Number of
measurements

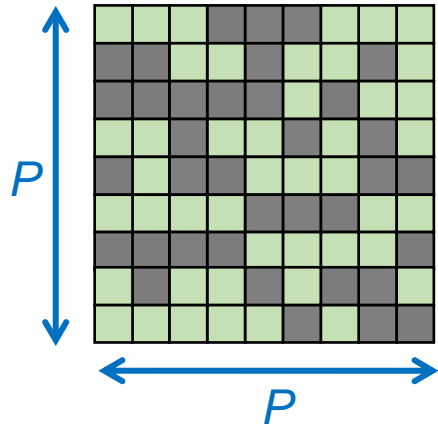
Image mean flux
(only depends on the scene)

* Assumes:

- Poisson-Gaussian noise [EMVA standard, 2021]
- Hadamard = negative part/S matrix
- Pseudoinverse reconstruction

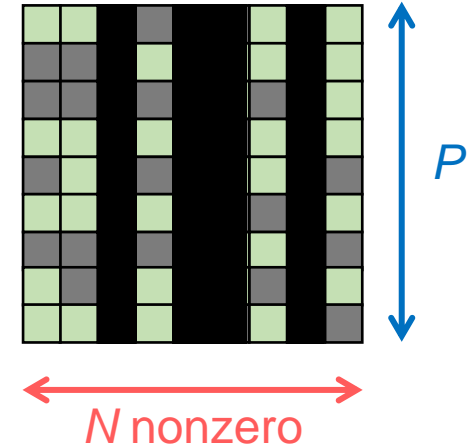
$$\text{Mse}(\mathbf{f}^*) \approx \frac{1}{t} [2Q\bar{f} + 4Mf_{\text{ref}}], \quad M \geq Q$$

Full



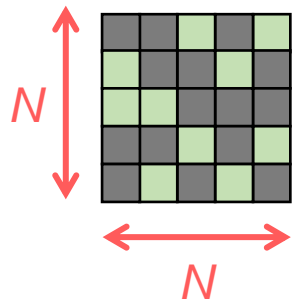
$$\text{Mse}(\mathbf{f}^*) \approx \frac{1}{t} [2P\bar{f} + 4Pf_{\text{ref}}],$$

Masked



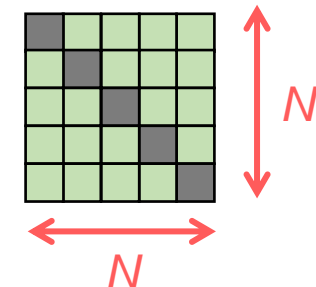
$$\text{Mse}(\mathbf{f}^*) \approx \frac{1}{t} [2N\bar{f} + 4Pf_{\text{ref}}],$$

Non imaging

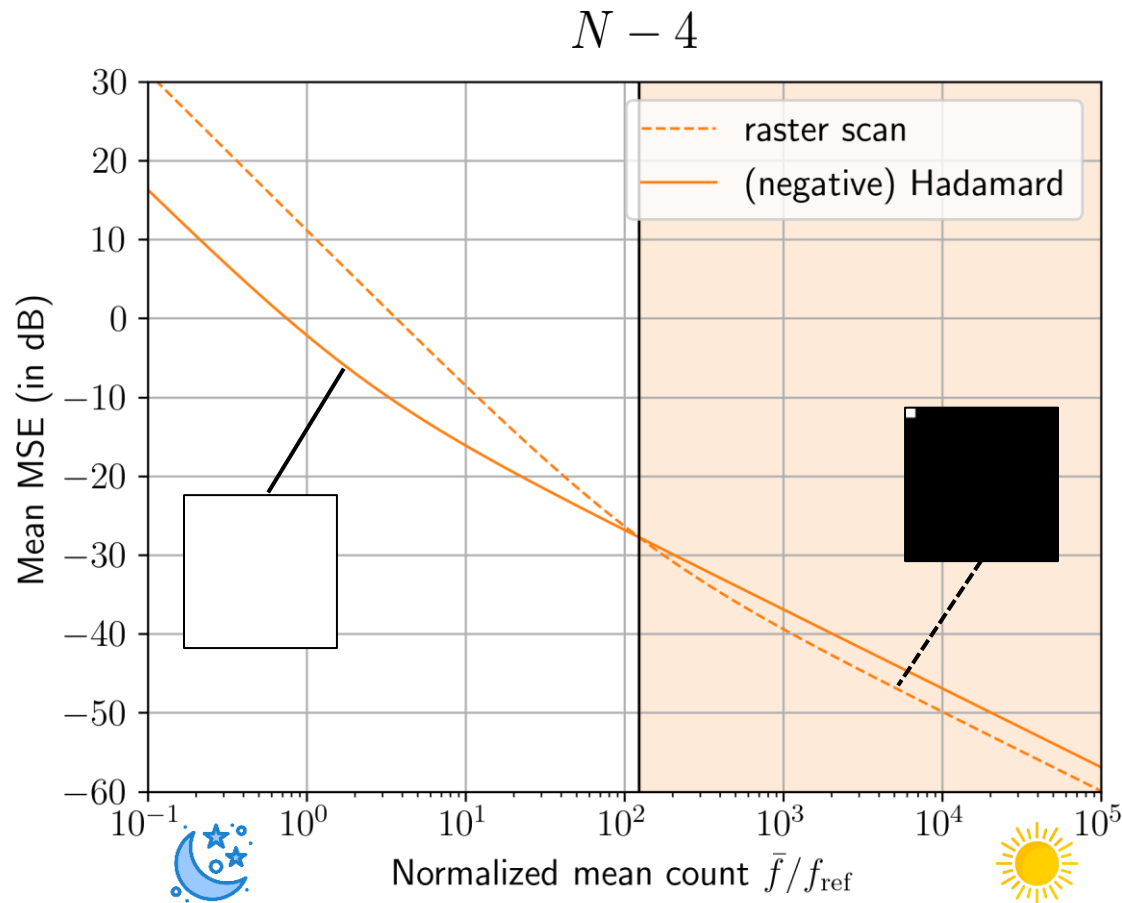


$$\text{Mse}(\mathbf{f}^*) \approx \frac{1}{t} [2N\bar{f} + 4Nf_{\text{ref}}],$$

Raster Scan

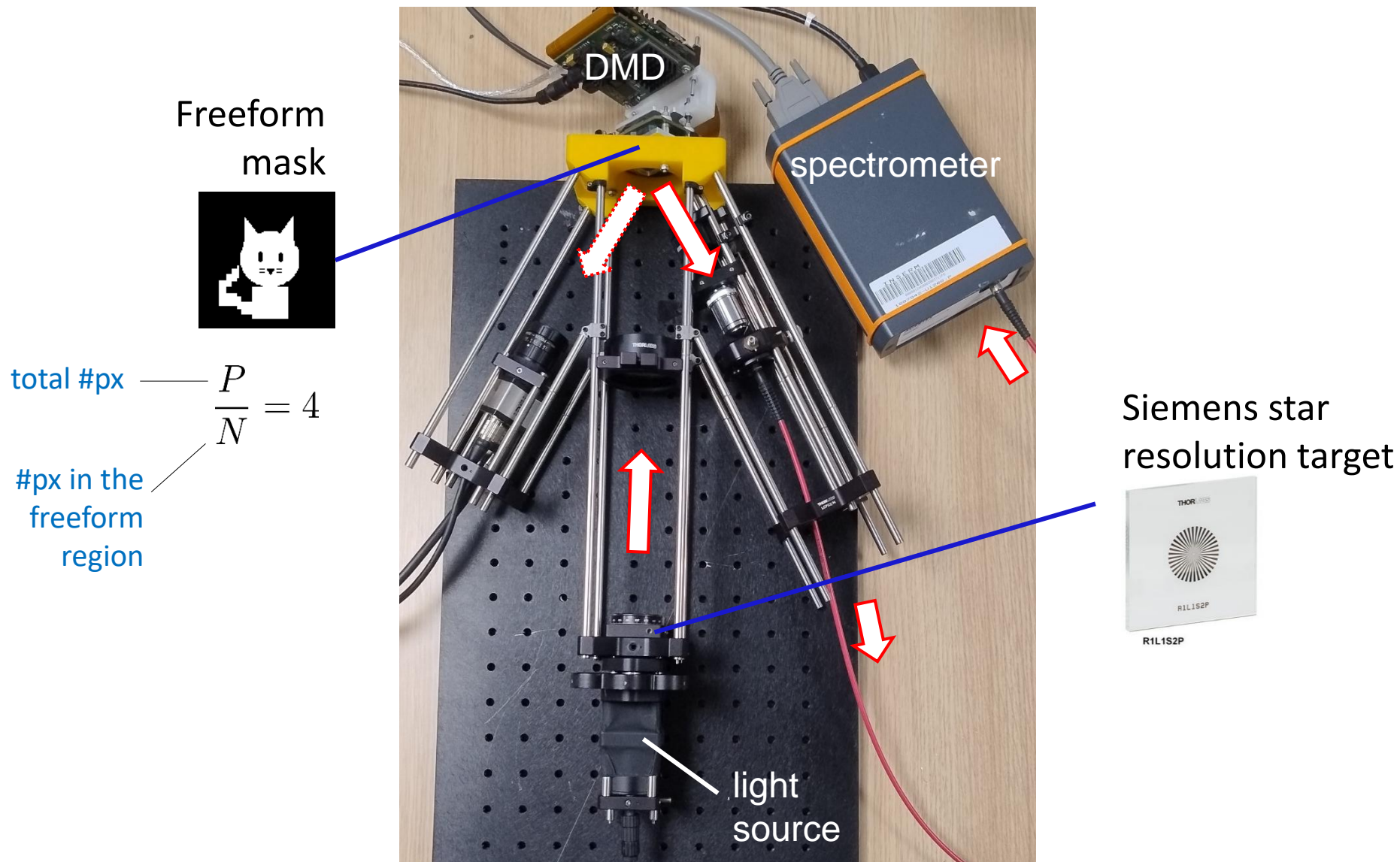


$$\text{Mse}(\mathbf{f}^*) \approx \frac{1}{t} [N\bar{f} + N^2f_{\text{ref}}],$$



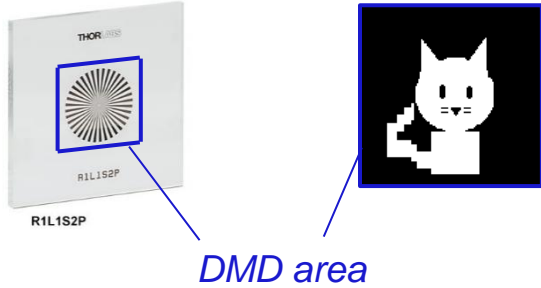
→ **Hadamard** multiplexing is more effective in **low-light** conditions.

4.
Experimental
Results

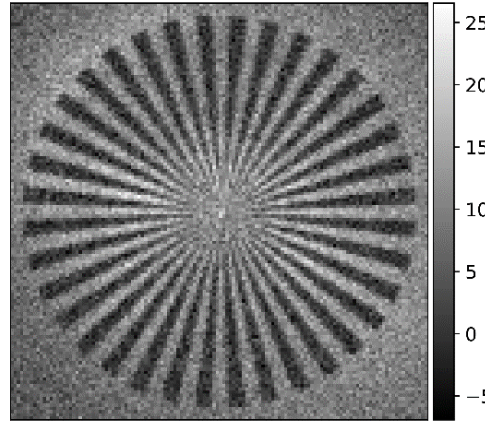


Siemens star resolution target

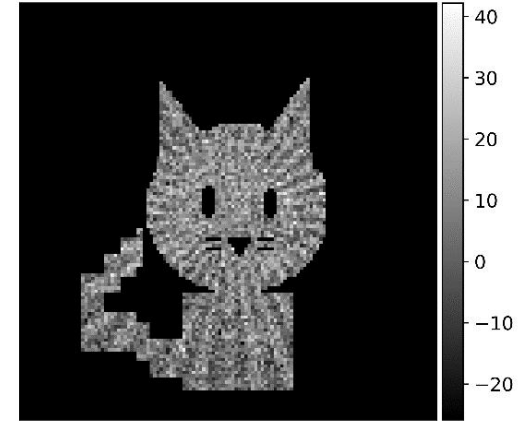
Freeform region



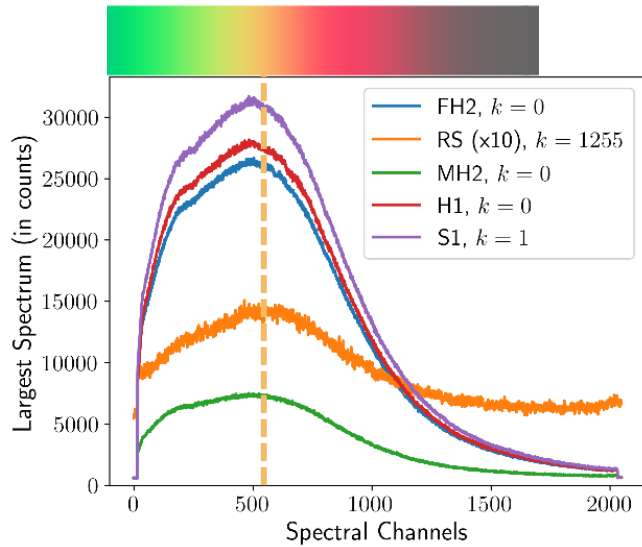
Full



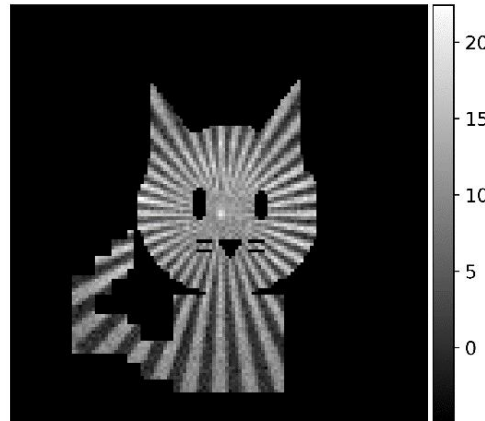
Raster Scan



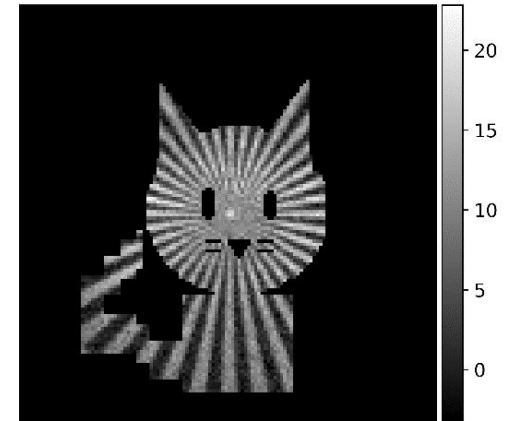
581 nm



Masked



Non imaging



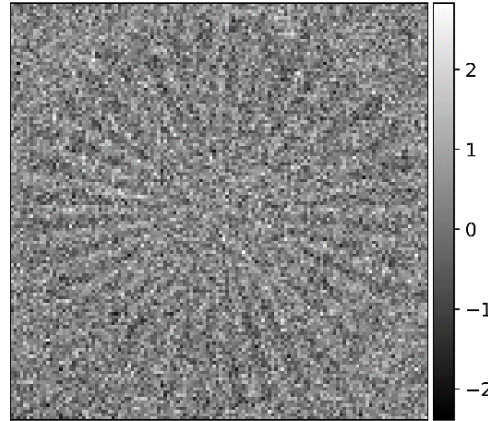
Siemens star
resolution target

Freeform
region

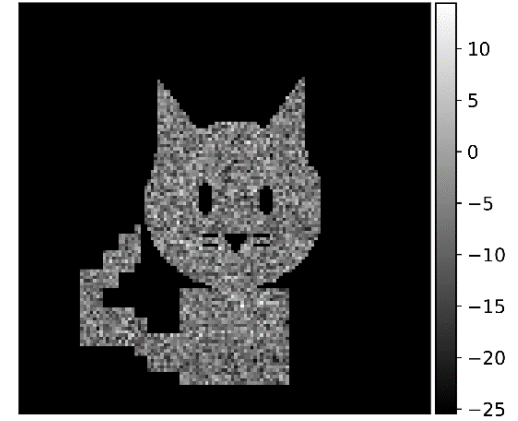


DMD area

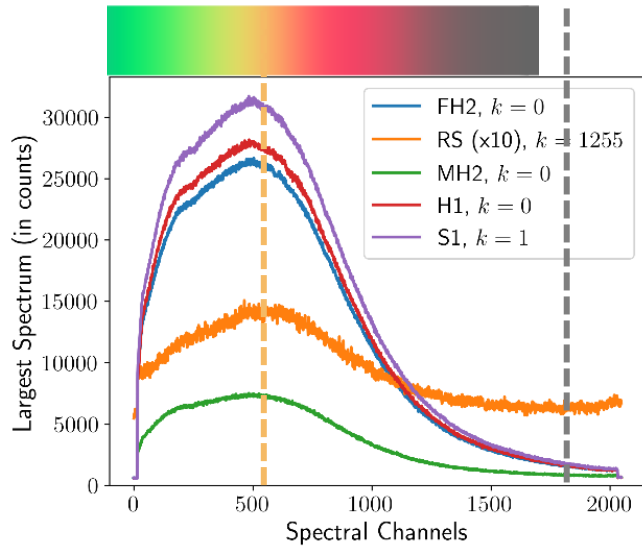
Full



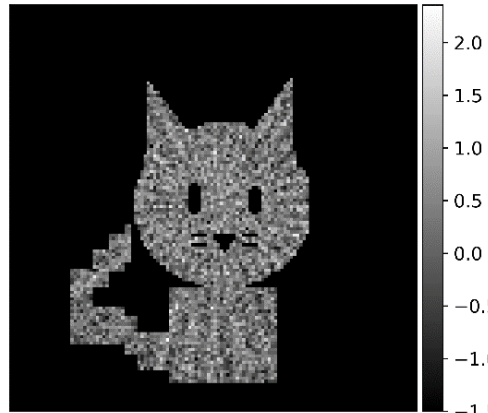
Raster Scan



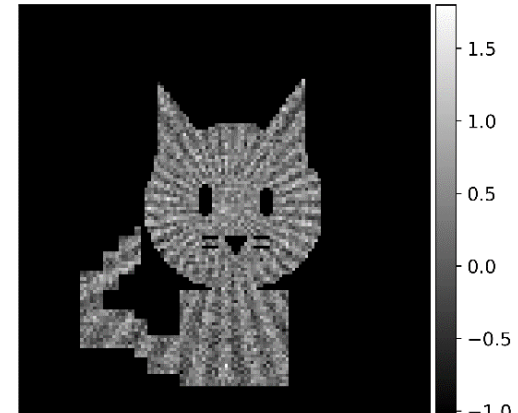
726 nm

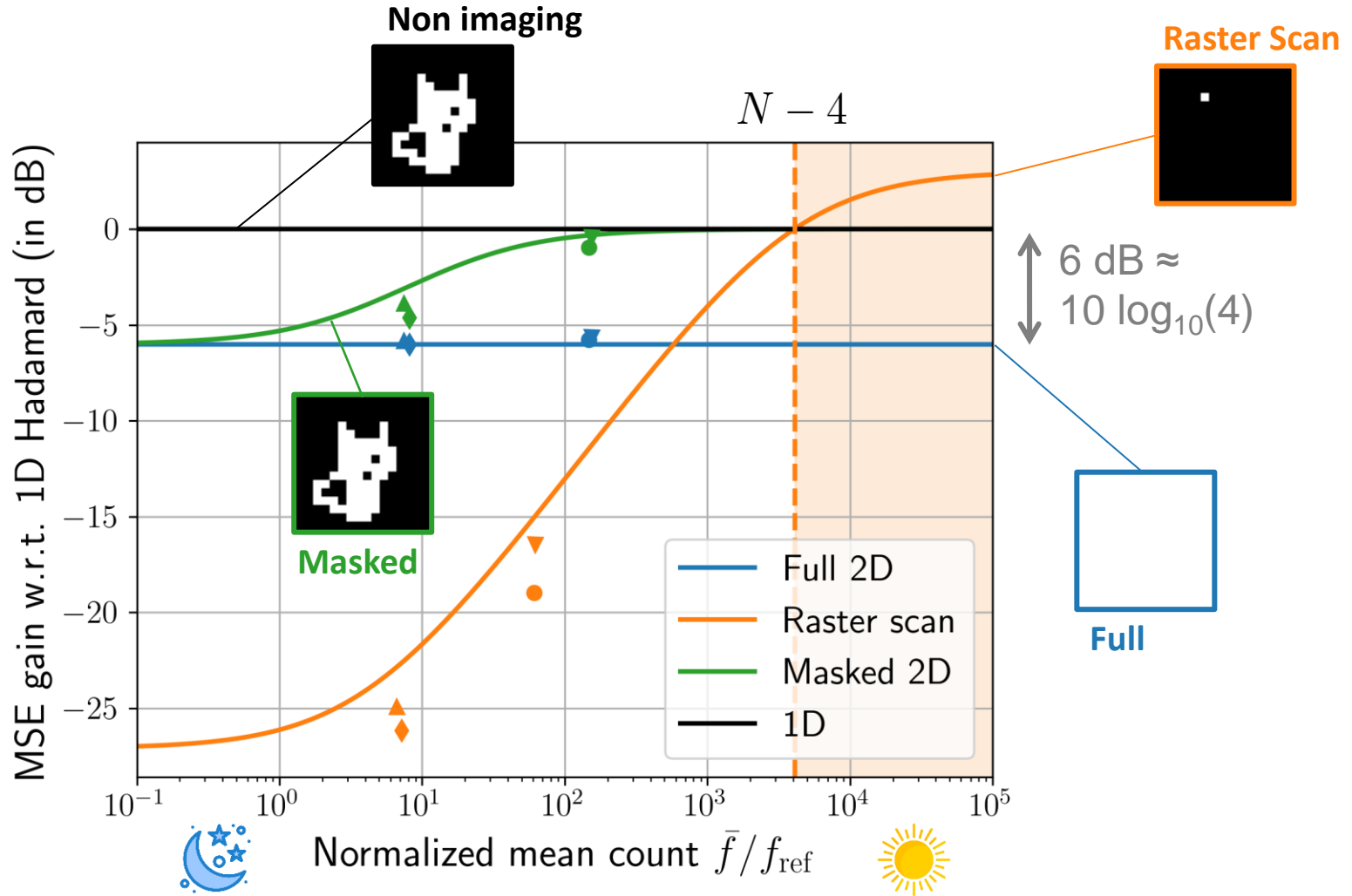


Masked



Non imaging





Note: The figure above was obtained using a split Hadamard matrix, not a negative Hadamard matrix, as the acquisition matrix.

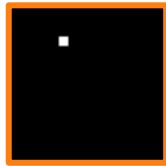
1. Freeform imaging



Non Imaging



Masked



Raster Scan

→ The 'non imaging' Hadamard matrix should be preferred to the masked one

2. We have characterized the MSE of Hadamard freeform imaging

→ **Freeform** imaging **outperforms** raster scan in **low/intermediate light** conditions

3. We have validated our results **experimentally** in hyperspectral imaging in the visible range

